LIGO Signals from the Mirror World



1. PARITY VIOLATION MIRROR AND PARTICLES

of the Talk

Plan 2. GRAVITATIONAL WAVES: THEORY **AND DETECTION**

> **3. UNEXPECTED LIGO EVENTS**

4. WORLD IS RULED **BY DARK FORCES**

Parity Violation Mirror Symmetry CP-violation and Mirror Symmetry CP-violation Mirror Partners Mirror World If our Universe = SM + SM' **Motivations Shadow Particles Theory of GWs GWs Detection Efforts** What Might Make GWs **LIGO Observatory Present Situation Events Detected So Far Masses in the Stellar Graveyard Unexpected Events Hierarchical Mergers GW190425: Most Massive BNS 3.** 4^{+0.3}₋₀₁ M_☉ GW190814 & GW200210_092254: BH Mass-gap **Binary Objects Creation Mechanisms Theoretical BBH Merger Rate Problems with LIGO Data** In Mirror World **Star Formation Rate** LIGO Signals from Mirror World

1. PARITY VIOLATION AND MIRROR PARTICLES

Parity Violation

 $\tau - \theta$ puzzle: Two different decays were found for charged strange mesons:

 $heta^+ o \pi^+ \pi^0$ $\tau^+ \rightarrow \pi^+ \pi^+ \pi^-$ parity = -1

parity = +1

Both reactions represent decay of the same particle, K^+ , with violation of parity.



"I cannot believe God is a weak left-hander" Wolfgang Pauli



Mirror Symmetry





C.N. Yang



Parity (**P**) violation means non-equivalence of **Left**- and **Right**-handed coordinate systems. To restore the symmetry, reflection should be generalized. Together with **P**transformations, particles should be changed by their mirror partners (*Lee*, *Yang* - *1956*).

CP-violation and Mirror Symmetry



Lee, Landau,... (1957) an economic solution: CP-invariance assumes that antiparticles play the role of mirror partners. Then non-conservation of parity can be introduced without assuming asymmetry of space with respect to inversion.

Dirac theory is symmetric under CP transformations!





CP-violation



RAFP Tbbilisi - 29 Sep 2022

Anti-world



The equivalence between L and R systems is restored if a hidden mirror sector exists in which parity is violated in the opposite way and reflection is accompanied by change of ordinary particles by their mirror partners.

Mirror partners are strictly symmetric to ordinary particles. Therefore they can not have ordinary electromagnetic and strong interactions (doubling of atomic levels, or pion states). Successive analysis has shown that (O) and (M) also can not share W and Z boson mediated weak interaction.



Mirror matter, also called **Shadow matter** or **Alice matter**, is a hypothetical counterpart to ordinary matter. One needs to suppose that there is no (or very weak) common interaction between ordinary particles and their mirror partners, except of gravity. Initially it was assumed that all initial conditions, masses and coupling constants of mirror particles were strictly symmetric to ordinary ones.



 $G = G_{SM} \times G'_{SM} = [SU(3)_c \times SU(2)_L \times U(1)_Y] \times [SU(3)_c' \times SU(2)_R' \times U(1)_Y']$ $L = L_{SM} + L'_{SM} + L_{int} \text{ is symmetric under parity exchange: } L_{SM} \leftrightarrow L'_{SM}$



The **Pati-Salam** model (the first **GUT - 1974**) is **L** – **R** symmetric. Predicts a high energy **R**-handed weak interaction with heavy **W'** and **Z'** bosons and **R**-neutrinos. **Gauge symmetry group:** $(SU(4) \times SU(2)_L \times SU(2)_R)/Z_2$



Motivations

Elegance suggests exact parity conservation → MIRROR MATTER

Historic precedence:

Lorentz invariance ↔ antiparticles ✓ Exact parity symmetry ↔ mirror particles ?

Mirror matter can explain several physical phenomena: *In cosmology:*

- Viable DM Candidate;
- High energy cosmic events.

In astrophysics:

- MACHOs;
- Isolated planets and stars;
- Drag force on Pioneer 10, 11;
- Missing comets;
- Mirror bodies bombardment on Earth.

In particle physics:

- Higgs mirror Higgs mixings;
- Neutrino sterile neutrino mixings;
- Photon-mirror photon kinetic mixing;
- Anomalous disappearance of neutrons.

Shadow Particles

Properties:

- > Feebly coupled to SM particles and have the same masses;
- > Interact with each other with shadow (SM') interactions;
- Stable, singlets under SM;
- > Scarce at the scale of the solar system but could dominate on a larger scale.

Can interact with us via:

Gravity;

Arkani-Hamed, Dimopoulos, Dvali, Kaloper, JHEP 0012 (2000) 010

□ Mixing with ordinary particles in the neutral sectors.

Problems:

- Symmetric initial conditions equal *M* and *O* baryonic densities;
- Mirror particles double species in BBN He abundance > 28%.

Solution:

The temperature of the mirror sector is a few times smaller!

2. GRAVITATIONAL WAVES: THEORY AND DETECTION

R -= 19, R=- 2

A CASSED

Theory of GWs

Gravitational Wave (GW) solution of linearized Einstein equations correspond to space-time fluctuations that propagate through the Universe.

The main stages in development of the **GW** theory are:

□ In **1916** Albert Einstein claimed that gravitational waves were an inevitable consequence of the linearized General Relativity.

□ In **1936 Einstein**, together with **Nathan Rosen**, derived the opposite result and submitted the paper "*Do gravitational waves exist*?" to *Physical Review*.

□ On semi-centennial conference in **Bern** (1955) **Rosen** showed that the energy-momentum pseudo-tensor for **Einstein-Rosen** waves is vanishing.

On Chapel Hill conference (1956) Richard Feynman proposed a "sticky-bead" experiment to show that GWs carry energy.



GWs Detection Efforts

The main stages in the **GW** detection are:

- □ First attempts were made in the **1960**s by **Weber** by means of aluminum resonant bar detectors;
- □ In **1963 Gertsenshtein** and **Pustovoit** suggested using **Michelson-Morley** type interferometers;
- □ In 1974 Taylor and Hulse detected the binary pulsar PSR B1913+16. Later, some more relativistic binary pulsars were discovered. Orbital period decay in all cases is consistent with the energy loss predicted by General Relativity. Nobel Prize for Physics in 1993 for the first observational evidence for GW;
- On 11 Feb. 2016, the LIGO/Virgo collaboration announced the first direct observation of GWs in Sep. 2015. The 2017 Nobel Prize in Physics Barish. Thorne and Weiss:
- □ The space missions: LISA (ESA, to launch in 2037?) and DECIGO (Japan, to launch in 2027?) were proposed.



What Might Make GWs







Neutron star spins

at frequency

Cosmological Signal: "stochastic background"

SASKATOON

3 YEAR DATA COBE DMR **4 YEAR DATA**

LIGO Observatory



LIGO - gravitational wave detectors in Hanford and Livingston, 4 km tunnels, separation 3,000 km. VIRGO - in **Cascina**, Italy

New Era of Multi-Messenger Astrophysics!

"Observation of Gravitational Waves from a Binary Black Hole Merger" LIGO/ Virgo, Phys. Rev. Lett. **116**, 061102 (11 February 2016)

"GW170817: Observation of Gravitational Waves from a Binary Neutron Star Inspiral"

LIGO/Virgo, Phys. Rev. Lett. 119, 161101 (16 October 2017)







Present Situation

Intensive efforts to find burst, binary coalescence, continuous wave, stochastic **GW**s in coincidence with **EM** and neutrino signals; **Black Holes of Known Mass**

> Stochastic background of relict GWs 1 = 1 could be observed in a few years;

> > launch 2037 (earlier launch 2034?);

The universe has more stellar mass **BH**s than expected;





First KAGRA detections with LIGO/Virgo were reported on 11 November 2021. The start of O4 Observing Run is scheduled for March 2023.



Mercury

Earth

Venus

relative orbit

of spacecraft

3rd generation detectors: **Einstein Telescope** (EU) and Cosmic Explorer (US), with atom interferometers in 0.1-10Hz (between LISA and LIGO-Virgo).

Events Detected So Far

After the analysis of first three observing runs

O1, O2, O3a & O3b

there are 90 events with a probability of

 $P_{astro} > 0.5$

being of astrophysical origin.

"GWTC-3: Compact Binary Coalescences Observed by LIGO and Virgo During the Second Part of the Third Observing Run" LIGO & VIRGO & KAGRA, arXiv: 2111.03606

Merger objects:	BH-BH	NS-NS	BH-NS	BH-Mass gap
Number of events:	84	2	2	2

Only one NS-NS merger had an accompanying Electromagnetic counterpart!

3. UNEXPLANMED LIGO EVENTS

Masses in the Stellar Graveyard

LIGO-Virgo-KAGRA Black Holes LIGO-Virgo-KAGRA Neutron Stars EM Black Holes EM Neutron Stars



BH-lower mass gap Upper mass gap, intermediate mass **BH**

NS-NS

Interactive figure: <u>https://ligo.northwestern.</u> <u>edu/media/mass-</u> <u>plot/index.html</u>

Unexpected Events

GW190521 & GW190426_190642: First ever observation of intermediate mass BHs

 $95M_{\odot} - 69M_{\odot} \rightarrow 156M_{\odot}$ $107M_{\odot} - 77M_{\odot} \rightarrow 175M_{\odot}$

Many models of star evolution predict existence of upper mass gap $65M_{\odot} - 135M_{\odot}$ for remnant compact objects.

<i>He</i> core mass	Process	Remnant compact object	
32 − 65 <i>M</i> ⊙	Pulsational Pair-instability	≲ 65 <i>M</i> ⊙	
65 — 135 M _☉	Pair-instability (e+e- pairs)	Explodes – no remnant	
≳ 135 <i>M</i> ⊙	Direct collapse into BH	≳ 135 <i>M</i> ⊙	

Hierarchical Mergers

Merger rate of **GW190521**-like events: $f_{exp} = 0.02 - 0.43 \text{ Gp}c^{-3}yr^{-1}$ LIGO /Virgo, ApJ. Lett. 900 (2020) L13 Theoretical estimate: 1g BH $f = f_{1g} \times f_{triple} \times f_{survival} \times f_{merger}$ Liu & Lai, MNRAS 502 (2021) 2049 $f_{1a} \sim (10 - 100) \, \text{Gp} c^{-3} y r^{-1}$ 1g+1g event LIGO /Virgo, Phys. Rev. X 9 (2019) 031040



Gerosa. & Berti, Phys. Rev. D 95 (2017)

 $f_{triple} \simeq 50 \%$ $f_{survival} \simeq 60 \%$ $f_{merger} \simeq 20 \%$ $f_{theor} = 0.6 - 6 \, \text{Gpc}^{-3} yr^{-1}$ In price of extremal assumptions!

GW190425: Most Massive BNS $3.4^{+0.3}_{-0.1} M_{\odot}$

X-ray binaries (normal star and a collapsed star) give the lower mass gap $2-5M_{\odot}$ for BHs!



GW190425-like systems could be obtained as a result of evolution of ultra-tight binary He-star – NS systems;

LIGO/Virgo, ApJ Lett. **892** (2020) L3

A phase of mass transfer from a post-helium main-sequence star onto NS is required.

GW190814&GW200210_092254:BH Mass-gap



First components are clearly **BH**s.

Originsofsecondcomponentswith masses $2.59^{+0.08}_{-0.09}M_{\odot} \& 2.83^{+0.48}_{-0.43}M_{\odot}$ are controversial.

They are heavier than any known pulsars, and lighter than any known BHs so far.

4. WORLD IS RULED BY DARK FORCES

Binary Objects Creation Mechanisms

- Primordial Black Holes (Sasaki, Suyama, Tanaka & Yokoyama, 2018)
 PBH abundance is constrained by microlensing, CMB spectral distortion and wide binaries.
- 2. Astrophysical binary systems:
 - Common envelope evolution; (*Giacobbo & Mapelli, 2018*)
 - > Chemically homogenous evolution; (*Mandel & de Mink. 2016*)
 - > Dynamical processes in dense stellar clusters. (Askar, et al. 2017)

Main formation mechanisms predict low binary merging rates:

$$\mathcal{R}_{theor}^{BBH} \simeq 5 - 10 \, \text{Gp}c^{-3}yr^{-1} < \mathcal{R}_{LIGO}^{BBH} = 17 - 45 \, \text{Gp}c^{-3}yr^{-1}$$

LIGO &VIRGO & KAGRA, arXiv: 2111.03634

Theoretical BBH Merger Rate $\mathcal{R} = \frac{1}{2} \varepsilon P(\tau) N_{BH}$ $\varepsilon \simeq 0.01 - 0.001$ - Efficiency coefficient $P(\tau)$ - Delay time distribution (time to merging) Number of Black Holes: (Elbert, Bullock & Kaplinghat, 2018) $N_{BH} = SFR(z) \times \int \varphi(m)N(m) \int f(Z,m) \int \xi(M) dM dZ dm$ $\varphi(m)$ - Galactic mass function $\xi(M)$ - Initial mass function $N(m) = \frac{m}{\int M\xi(M) dM}$ - Number of stars in the galaxy of mass m f(Z, m) - Metallicity distribution function Star formation rate: (Madau & Dickinson, 2014) $SFR(z) = 0.015 \frac{(1+z)^{2.7}}{1 + [(1+z)/2.9]^{5.6}} M_{\odot} Mpc^{-3}yr^{-1}$ Peaks at: $z \sim 2 \approx t_{lookback} \sim 10.3$ Gyr

Problems with LIGO Data

- Observed Merger Rates are higher than theoretical predictions
- > Only 1 out of 90 events had EM counterpart, while:
 - BNS merger must always be accompanied by Gamma-Ray Bursts;
 - BH-NS mergers in many configurations should emit EM-radiation;
 - If **BH**s accrete matter they can also emit **EM-radiation**.
- Mass gap events observed

Suggestion:

GWs detected by **LIGO** may be emitted by the **Mirror World** binaries!

In Mirror World

M-world, along with *O*-world, was created by **Big Bang**, but with lower reheating *T*.

- > Constrain from Big Bang Nucleosynthesis: $x \equiv T'/_T < 0.64$
- > Certain leptogenesis mechanism gives: $1 \le \frac{n'_b}{n_b} \le 10$
- > Mirror World can explain all Dark Matter: $\frac{\Omega'_b}{\Omega_L} \approx 5$
- ➢ Helium abundance is higher: He 75-80 %
- > Stars are composed mostly of He, are more massive and evolve faster, e.g. $10M_{\odot}$ star with 75% He evolves ~ 10 times faster than normal star (24% He).
- > High He abundance increases Initial Mass Function: $\xi(M)' \sim 1.5 \times \xi(M)$
- > Star formation peaks at $z \sim 10 \approx t_{lookback} \sim 13.3$ Gyr: $SFR'(z) \sim 2.3 \times SFR(z)$
- > Number of stars: $N'(m) \sim 5 \times N(m)$
- > Number of black holes: $N'_{BH} \sim 10 \times N_{BH}$

For a review on mirror world see **Z. Berezhiani, 2005**

Star Formation Rate

In the period 0 < z < 14 more stars are formed in Mirror World relative to our:

$$\frac{\int_{0}^{14} SFR' \, dz}{\int_{0}^{14} SFR \, dz} = 2.3$$

The number of Mirror black holes $N'_{BH} \sim 10 N_{BH}$

Even if mirror matter does not make up all dark matter, or if formation of binary systems is not so efficient, the amplification factor still can be ~ 5 .



SFR in the interval 0 < z < 14In red: Mirror SFR' for different temperatures x = T'/T. (Madau & Dickinson 2014)

LIGO Signals from Mirror World

> Merging Rate with combining Mirror World amplification factors: $\mathcal{R}_{mirror} \sim 10 \times \mathcal{R}_{theor} \sim 50 - 100 \, \text{Gp}c^{-3}yr^{-1}$

coincides with **LIGO**'s bounds even if some assumptions of binary formation are relaxed;

Hierarchical mergers are more probable in Mirror World and merger rates of upper mass gap systems (GW190521 & GW190426_190642) would agree better even with less strict assumptions;

Production of 'heavy NSs' (GW190425) or lower mass gap objects (GW190814 & GW200210_092254) are easier in Mirror World, as they are dominated by He.

Conclusions:

In the Mirror World scenario:

- Number of binary systems is higher;
- BBH merger rate is amplified, coinciding better with LIGO estimations;
- Mass gap events could be better explained;
- Non-detection of EM-radiation is natural, since Mirror photons DO NOT interact with Ordinary particles.

Prediction: Binary compact objects' merger rates are order of 10 higher than expected and only 1 of 10 NS-NS events discovered by GW detectors may have EM-counterpart.

References:

- 1. "Gravitational Waves from Mirror World", MDPI Physics 1 (2019) 67;
- 2. "LIGO Signals from the Mirror World", MNRAS 487 (2019) 650;
- 3. "Binary Neutron Star Mergers with Missing Electromagnetic Counterparts as Manifestations of Mirror World", Phys. Lett. B 804 (2020) 135402;
- 4. "Unexpected LIGO events and the Mirror World", MNRAS 503 (2021) 2882.



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